OZONE FORMATION AND ION CHEMISTRY IN OXYGEN RADIOLYSIS

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Technical Report

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19. ABSTRACT (Continue on reverse if necessary and identify by block number) Detailed electron-impact and chemistry calculations have been carried out for					
situations corresponding to a	cnemistry calcu number of publ	iations nave ished oxvaer	radiolysi	ied out for s experimer	its.
The results fit the measured	ozone productio	n values wit	thin 25 per	cent, provi	ded
that the neutralization reaction of $0\frac{1}{4}$ with $0\frac{1}{2}$ or $0\frac{1}{4}$ yields two oxygen atoms, while the neutralization of $0\frac{1}{4}$ or $0\frac{1}{5}$ by $0\frac{1}{3}$ yields no odd oxygen species (0 or $0\frac{1}{3}$)					
except for that already present in the ions.					
In addition, to explain the steady-state ozone concentration attained by long- duration irradiation, the reaction $0.5 \pm 0.5 \pm 0.06$ is required, with a rate					
duration irradiation, the reaction $0\bar{3}+0_3\to 0\bar{2}+2$ 0_2 is required, with a rate coefficient within a factor of two of 5 x 10^{-13} cm ³ /s at room temperature and					
probably a temperature dependence of about exp (-1000/T).					
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CONVERSION TABLE

Conversion factors for U.5. customary to metric (S1) units of measurement.

(5ymbols of \$1 units given in parentheses)

To convert from	to	Multiply by
angstrom	meters (m)	1.000 000 X E -10
atmosphere (normal)	kilo pascal (kPa)	1.013 25 X E +2
bar	kilo pascal (kPa)	1.000 000 X E +2
barn	meter ² (m ²)	1.000 000 X E -28
British thermal unit (thermochemical)	joule (J)	1.054 350 X E +3
calorie (thermochemical)	joule (J)	4.184 000
cal (thermochemical)/cm ²	mega joule/m² (MJ/m²)	4.184 000 X E -2
curte	giga becquerel (GBq)*	3.700 000 X E +1
degree (angle)	radian (rad)	1.745 329 X E -2
degree Fahrenheit	degree kelvin (K) t _v	$=(t_{\bullet F} + 459.67)/1.8$
electron volt	joule (J)	1.602 19 X E -19
erg	joule (J)	1.000 000 X E -7
erg/second	watt (W)	1.000 000 X E -7
foot	meter (m)	3.048 000 X E -1
foot-pound-force	joule (J)	1.355 818
gallon (U.S. liquid)	meter ³ (m ³)	3.785 412 X E -3
inch	meter (m)	2.540 000 X E -2
jerk	joule (J)	1.000 000 X E +9
<pre>joule/kilogram (J/kg)(radiation dose absorbed)</pre>	Gray (Gy)**	1.000 000
kilotons	terajoules	4.183
kip (1000 lbf)	newton (N)	4.448 222 X E +3
kip/inch² (ksi)	kilo pascal (kPa)	6.894 757 X E +3
ktap	newton-second/m ² (N-s/m ²)	1.000 000 X E +2
micron	meter (m)	1.000 000 X E -6
mil	meter (m)	2.540 000 X E -5
mile (international)	meter (m)	1.609 344 X E +3
ounce	kilogram (kg)	2.834 952 X E -2
pound-force (1bf avoirdupois)	newton (N)	4.448 222
pound-force inch	newton-meter (N-m)	1.129 848 X E -1
pound-force/inch	newton/meter (N/m)	1.751 268 X E +2
pound-force/foot ²	kilo pascal (kPa)	4.788 026 X E -2
pound-force/inch ² (psi)	kilo pascal (kPa)	6.894 757
pound-mass (1bm avoirdupois)	kilogram (kg)	4.535 924 X E -1
pound-mass-foot ² (moment of inertia)	kilogram-meter² (kg·m²)	4.214 011 X E -2
pound-mass/foot3	kilogram/meter (kg/m³)	1.601 846 X E +1
rad (radiation dose absorbed)	Gray (Gy)**	1.000 000 X E -2
roentgen	coulomb/kilogram (C/kg)	2.579 760 X E -4
shake	second (s)	1.000 000 X E -8
slug	kilogram (kg)	1.459 390 X E +1
torr (mm Hg, O° C)	kilo pascal (kPa)	1.333 22 X E -1
(besser (w.a)	1.000 EF W F -1

^{*}THE BECQUEREL (BQ) IS THE SI UNIT OF RADIOACTIVITY; 1 BQ = 1 EVENT/S. **THE GRAY (GY) IS THE SI UNIT OF ABSORBED RADIATION.

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SECTION 1

INTRODUCTION

A high altitude nuclear explosion exposes the upper atmosphere to fast photoelectrons produced by absorption of the x-rays from the bomb, and to beta rays (megavolt electrons) from the radioactive debris. These electrons, and the secondary electrons they produce, directly dissociate some of the oxygen in the air, and also create ions and other reactive species that lead to further dissociation. The atomic oxygen thus created is of practical importance because it detaches electrons from negative ions, which increases the absorption of radio waves. In addition, it combines with molecular oxygen to form ozone (0_3) , which is a strong infrared emitter, particularly in the 9 to 11 μ m spectral region.

Current predictions of the amounts of atomic oxygen and ozone produced per unit energy deposited in the upper atmosphere are based on rather complex calculations of the relevant electronimpact and chemical processes, since no direct measurements are available. However, several laboratory measurements of the ozone production in irradiated air and oxygen at pressures typical of the lower atmosphere have been carried out and published. It is useful to utilize these experimental results to check or improve the complex air irradiation calculations, especially since the calculations involve several uncertain cross sections and rate constants. A logical procedure is to treat the pure oxygen case first, without the complication of nitrogen or water vapor reactions. This is done in the present analysis; even without the other air components it yields some significant results pertinent to atmospheric situations. A later report will treat the normal air case.

A few previous researchers have made limited comparisons of oxygen radiolysis measurements with chemical calculations. The most recent and most complete of these is that published by Willis et al. (Ref. 1) in 1970. Although these investigators were able to obtain reasonable agreement of calculations with measurements, they assumed rather than calculated the initial number of oxygen atoms produced by the irradiation. Moreover, although they used the best chemical rate coefficients available at that time, some of their values are now known to be inaccurate, and they neglected excited state reactions. Accordingly, a new and more complete analysis is needed in order to confirm or correct the conclusions of Ref. 1.

SECTION 2

INITIAL SPECIES PRODUCTION IN THE RADIOLYSIS OF OXYGEN

2.1 COMPUTATIONAL PROCEDURE.

The first step in making ozone calculations for comparison with the measurements is to determine the species produced initially by the electron bombardment. This requires a complex calculation of the excitation, dissociation and ionization processes that occur as the primary and secondary electrons slow down to thermal energies. Several groups of researchers have developed computer codes for making such calculations for any gas of interest, provided that the relevant electron-impact cross sections are input. Two groups have published a few results for pure oxygen (Refs. 2,3). However, neither group has given a set of results complete enough for the present purposes, nor did either employ the best available input cross sections.

Accordingly, new calculations of the initial species production in oxygen have been made at RDA. The results, are summarized in Table 1. The RDA calculations employed a "discrete energy bin" code, DPOSIT, which is based on a code originally developed by Dalgarno, Lejeune and Victor (see Ref. 4), modified by Stephens and Klein to handle higher energy primaries (Ref. 5), and further modified for greater input and output flexibility by the present writers and coworkers at RDA. It has been checked against a comparable code developed by Green and colleagues (Ref. 2) by applying it to irradiated air, and found to give results agreeing within 2 percent in every detail, when the same input cross sections are used.

Table 1. Initial production of atoms, ions and excited states in bombarded oxygen*.

Species O2(a ¹ Ag)	G Value (particles/100 eV)
O ₂ (b ¹ Σ _g)	1.7 †
O(³ P)	4.0
O(1D)	3.0
02	2.25
0+	1.0
е	3.25

^{*} The results are essentially identical for bombardment by 1 keV and 10 keV primary electrons.

[†] These values, which were used in the present chemistry calculations, were computed with 0.15% CO₂ in the O₂. In pure oxygen the values would be slightly larger, 8.9 and 1.8, respectively, but this makes a negligible difference in the ozone chemistry.

2.2 IMPORTANCE OF ELECTRON-IMPACT CROSS SECTIONS.

The principal source of error in electron degradation calculations is the inaccuracy in the basic electron-impact cross sections. Using limited measurements and theory, Porter, Jackman and Green (Ref. 6) derived algebraic expressions for the electronic excitation cross sections of oxygen, and Jackson, Garvey and Green (Ref. 7) did similarly for the ionization cross sections. However, more recent measurements, as well as some older measurements which they overlooked, show that their expressions are frequently significantly in error, sometimes by more than a factor of two. For the present analysis, a careful search for and evaluation of the best available cross sections was made, and the results were input to the DPOSIT code as numerical tables, avoiding the approximation of algebraic fits.

A complete DPOSIT calculation requires inputs and provides outputs for a large number of excited and ionized states. For present purposes, the most important outputs are the total amounts of oxygen atoms and 0+ and 0½ ions produced. In addition, the low metastable states of 02 and 0 are involved in ozone-destroying reactions, and so should be included. Moreover, several other excited states also need to be included in the DPOSIT calculation, because excitation removes energy from the colliding electrons, leaving less energy for dissociation or ionization by subsequent collisions. In addition, dissociation of molecules by electron impact actually occurs by excitation of unstable electronic states which spontaneously dissociate; inclusion of such states is, of course, essential to the present analysis.

Since so much depends on input cross sections, the sources of the data used in the present calculations will be summarized below in some detail, starting with the lowest electronic states and working upward.

2.3 CROSS SECTIONS FOR NEUTRAL EXCITATION AND DISSOCIATION.

For the low 0_2 metastable $a^1\Delta$ and $b^1\Sigma$ states, smooth curves were drawn through the data of Linder and Schmidt (Ref. 8), Trajmar et al. (Refs. 9 and 10), and Wakiya (Ref. 11), which are generally consistent in their overlapping energy regions. For the combined A, C and c states near 5 eV, which are too close together to be resolved in electron-impact measurements, the data of Refs. 10 and 11 were used. Since no measurements below 15 eV incident energy are available, the cross section was assumed to fall rapidly below this energy, in analogy to the rapid decrease of the cross section for the neighboring B state, discussed below.

A curve for the important B3x state, which contributes over half of the total oxygen atom production, was drawn through the data of Trajmar et al. (Ref. 10) and Wakija (Ref. 12). A similar curve for the low Rydberg states between 9.7 and 12.1 eV was based on the data of Wakiya (Ref. 12). For both states the published data extend only down to incident energies of 20 eV. Below this energy both cross sections were assumed to decrease rapidly, in accordance with a preliminary unpublished value by Shyn and Sharp (Ref. 13) for the B state at 15 eV.

As mentioned earlier, electron-impact dissociation of molecules generally takes place by excitation to unstable electronic states which spontaneously dissociate. In oxygen

this is believed to occur for essentially every excitation to the A + C + c and B states, and to every Rydberg state below 12.1 eV (the ionization energy of 02). For the A + C + c and B states this conclusion follows from the Franck-Condon principle applied to the potential curves (Ref. 14) and also from the observed electron energy loss (Ref. 10-12) both of which indicate production of these states above their dissociation limits, so that the two oxygen atoms immediately fly apart. For the Rydberg states, the observed broadening of the ultraviolet absorption lines and the failure to detect corresponding emission lines indicate that these states "predissociate," i.e., spontaneously dissociate by interaction with neighboring dissociating states. sections for all of these dissociating states could be combined into a single "dissociation cross section." However, because the energy lost by the incident electron varies with the particular state excited (the excess over the dissociation energy going into kinetic or excitation energy of the atomic fragments), it is desirable in DPOSIT calculations to treat each dissociating state separately.

In addition to the states discussed above, Porter et al.

(Ref. 6) also considered a number of higher Rydberg states.

Since none of these states are observed in optical emission, they must either predissociate or autoionize. Porter et al. arbitrarily assumed a 50-50 split between these two processes for most of the states, while photoabsorption and photoionization data (Ref. 15) show that autoionization is generally favored. In addition, the estimates of Porter et al. for the cross sections of these states were based on their estimates of the cross sections for the excited states of 05

(the Rydberg "cores"), which they generally overestimated, as discussed below. Consequently, their Rydberg cross sections are probably too high.

It would be difficult to re-estimate cross sections for all of the 18 higher Rydberg states treated by Porter et al., and the results would still have considerable uncertainty. Fortunately, these states contribute less than 5 percent to the total dissociation cross section and less than 15 percent to the total ionization cross section. the ionization error can be largely removed by constraining the total of all the ionization cross sections to the measured total ionization value. Accordingly, for the present DPOSIT calculations only the two lowest (and most strongly excited) Rydberg states above 12.1 eV, i.e., the $(a^4\Pi)3R$ and $(A^2\Pi)3R$ states at about 12.7 and 13.3 eV, were included. The cross sections used were those Porter et al. (Ref. 6). Although, as discussed above, these cross sections are probably too large, this is offset by the neglect of higher Rydberg states. In accordance with photoabsorption and photoionization data (Ref. 15), these states were assumed to dissociate spontaneously 30 percent of the time and autoionize 70 percent of the time.

2.4 CROSS SECTIONS FOR IONIZATION.

The cross section for ionizing ${\rm O}_2$ to its ground ion state, ${\rm X}\,{\rm 2\Pi}$, was derived by subtracting the cross sections for Rydberg autoionization, just discussed, and for excited ion states, discussed below, from the total ionization cross section. The total cross section was taken from the measurements of Rapp and Englander-Golden (Ref. 16) up to 200 eV and of Schram et al. (Refs. 17, 18) above 500 eV, with a smooth

transition in the intermediate energy region where the data of the latter lie about 10 percent lower than those of the former.

Turner et al. (Ref. 19) measured the fraction of metastables in an 0^+_2 beam produced by bombarding 0_2 by electrons of energies up to 200 eV. This fraction rises from zero near 16 eV to a value of 0.33 above 70 eV. These results can be converted to a cross section for producing metastables by multiplying by the total oxygen ionization cross section, determined above, and by the fraction of the total ions $(0^+_2 + 0^+)$ that are 0^+_2 , measured by Rapp et al. (Ref. 20). The only known or predicted metastable state of 0^+_2 is the $a^4\Pi$ state at 16.11 eV above the ground state of 0_2 (Ref. 14). This 0_2^+ state can be produced either directly or by the rapid radiative decay of the $b^4\Sigma$ and $c^4\Sigma$ states. However, the c-state apparently contributes little to the metastable production, since it primarily dissociates (Ref. 6). Accordingly, the a-state cross section has been obtained by subtracting from the metastable cross section, the cross section for producing the b-state, derived below. It might be noted that this new a-state cross section is about two-thirds of that employed by previous investigators (Refs. 2, 3, 5-7).

The cross sections for exciting a few emission bands of the 0^+_2 $A^2\Pi$ state have been measured by Korol' et al. (Ref. 21), while a single point for one band at 100 eV has been given by McConkey and Woolsey (Ref. 22). However, the ratios of several band intensities given by Korol' et al. disagree seriously with the ratios obtained in radiative studies (Ref. 23). Moreover, the cross section of Korol' et al. falls off below 100 eV with unrealistic steepness. Fortunately, both this data and that of McConkey and Woolsey show that the A-state

cross section is less than one percent of the total ionization cross section, and so can be neglected for present purposes. (The correct cross section is clearly over an order of magnitude smaller than the semiempirical cross section derived in Ref. 7).

The cross section for exciting the 0^+_2 b $^4\Sigma^-_g$ state has been measured with modern methods by five different groups of investigators (Refs. 22, 24-27), omitting the discordant measurements of Korol' et al. (Ref. 21) for reasons discussed earlier. All obtained curves of essentially the same shape, with a maximum around 100 eV, but their absolute magnitudes are significantly different, presumably due to differences in normalization (often a problem in this field). After correction for incomplete coverage of all the b-state emission bands, the cross sections of three of the investigating groups (Ref. 22, 24, 25) agree within 10 percent, while that of Borst and Zipf (Ref. 26) is about 40 percent lower, and that of Srivastava (Ref. 27) is about 50 percent lower.

As discussed earlier, the sum of the a- and b-state cross sections is fixed by the metastable measurements of Turner et al. (Ref. 19). If the b-state cross section of Refs. 22, 24 or 25 is assumed correct, the two cross sections are found to be roughly equal, while if that of Borst and Zipf (Ref. 26) or Srivastava (Ref. 27) is used, the b-state cross section is found to be about 1/2 or 1/3 that of the a-state.

Guidance in choosing among these alternatives can be obtained from optical measurements. Theory and measurements (Ref. 28) show that above a few hundred eV the electron-impact cross section for ionization is approximately proportional to the photoionization cross section (or differential optical

f-number) integrated over photon energies with a weighting factor that emphasizes the lower energies. Samson et al. (Ref. 29) have measured the partial photoionization cross sections for producing the b-state and the combined a + A states (unresolved), between threshold and 41 eV. Over most of this region the a + A cross section is about twice that of the b-state; moreover, it is quite large in the 16-18 eV region, which is below the threshold of the b-state. as discussed above, optical emission data show that the contribution of the A state to the a + A combination is small, one concludes that the electron-impact cross section for producing the a-state is at least twice that for producing the b-state. This rules out the b-state cross sections of Refs. 22, 24 and 25. Of the two remaining sets of data, that of Borst and Zipf (Ref. 26) has been chosen, because they are more experienced workers and their data are smoother. (These values are only about half as large as the semiempirical values of Ref. 7.)

Electron impact on 0_2 also produces several excited states of 0_2^+ above the b-state. Apparently all of these states spontaneously dissociate into $0^+ + 0$ (Ref 6), except for a small fraction of the c-state, which can be neglected for present purposes. Hence, they are included in the dissociative ionization cross section measured by Rapp et al. (Ref. 20). Much less is known about the partition of this cross section among the different excited states, but fortunately the DPOSIT calculations are less sensitive to the partition than to the total. Jackman et al. (Ref. 7) used the limited experimental evidence to divide these states into three groups, with thresholds at 20, 23 and 37 eV. Their groups and cross sections have been used in the present analysis, except that some adjustments were made to the 20 eV cross section to make

the total of the three agree with the measurements of Rapp et al. (Ref. 20).

2.5 RESULTS OF ELECTRON IMPACT CALCULATIONS.

The results of the DPOSIT calculation are summarized in table 1, with the following modifications: All predissociating excited states (discussed above) are replaced by their dissociation products, 0 + 0 or $0^+ + 0$. The ground state (3P) and metastable (1D) atoms are tallied separately, which requires information about the electronic excitation of the dissociation products. The lowest dissociating states, A + C + c, and the dissociatively ionizing states below 20.7 eV, have insufficient energy to produce excited O atoms. The well-studied $^{\circ}_{2}$ B state is known to dissociate into $^{\circ}_{1}$ + $^{\circ}_{1}$ + $^{\circ}_{1}$ + $^{\circ}_{1}$ products of the higher excited states are less well established, but they contribute only about 20 percent of the total O atoms. Consequently, for these states the product assignments of Porter et al. (Ref. 6) have been used without further study. The small amount of $O(^{1}S)$ product they suggest is assumed here to be quenched immediately to $O(^{3}P)$.

The state of excitation of the ionized products, 0_2^+ and 0^+ , has been ignored for present purposes, since on collision with 0_2 , excited 0_2^+ is rapidly quenched and excited 0^+ rapidly transfers its charge (Ref. 30). Such reactions might possibly yield metastable 0_2 or 0. but since there is no evidence for this possibility, it has been neglected. The DPOSIT calculation gives $G(0_2^+) = 2.42$ and $G(0^+) = 1.04$, where the G values are defined as the number of particles produced per 100 eV of deposited energy. The sum of these values, 3.46, slightly exceeds the experimental value, G(ions) = 3.25, derived from several careful measurements in oxygen showing that 30.8 eV is

required to produce each ion pair (Ref. 31). Although the difference is only 6 percent, the experimental value appears accurate to about 1 percent. Accordingly, for table 1 the calculated values of $G(0_2^+)$ and $G(0^+)$ have been decreased by 6 percent to give agreement with the ionization measurements.

The DPOSIT calculation was carried out for 1 keV primary electrons, while the oxygen radiolysis experiments generally used bombardment energies of hundreds of keV. However, theoretical calculations (Ref. 6) and ionization measurements (Refs. 31, 32) indicate that the G values are constant within about 5 percent for primary energies above 1 keV.

SECTION 3

CHEMICAL REACTIONS IN OXYGEN RADIOLYSIS

3.1 SOURCES OF BASIC DATA.

The chemical reactions believed to be involved in oxygen radiolysis are listed in table 2, together with their rate coefficients. Much of the chemical kinetic data is taken from standard compilations of laboratory measurements (Refs. 33-38), and some from more specialized papers (Ref. 39-49). Several rate coefficients, however, have not been measured, so have had to be estimated by analogy with similar reactions or by use of indirect evidence, as indicated in the footnotes to the table.

3.2 ION-MOLECULE PROCESSES.

In oxygen radiolysis, the initial bombardment produces 0^+_2 , 0^+ , and thermal electrons. The electrons very rapidly attach to 0^-_2 :

$$e + o_2 + o_2 \rightarrow o_2^- + o_2^-.$$
 (1)

The primary 0+ ions also rapidly charge exchange with 02:

$$0^+ + 0_2 \rightarrow 0_2^+ + 0.$$
 (2)

The $0\frac{1}{2}$ and $0\frac{1}{2}$ ions then evolve by charge transfer, ion-neutral clustering, and other processes. The basic clustering reactions involve $0\frac{1}{2}$:

$$0_2^- + 0_2^- + 0_2^- + 0_4^- + 0_2^-$$
 (3)

$$o_2^+ + o_2 + o_2 \rightarrow o_4^+ + o_2$$
 (4)

Table 2. Reactions and Rate Coefficients for the Pure Oxygen System.

Reaction	Rate Coefficient (a)	Reference or Footnote
$e + 0^{+}_{2} \rightarrow 0 + 0(^{1}D)$	$2.1 \times 10^{-7} (\frac{300}{T})^{0.63}$	33
$0^+ + 0_2 \rightarrow 0_2^+ + 0$	1.9 X 10 ⁻¹¹	34
$0_2^+ + 0_2^- + 0_2^- \rightarrow 0_4^+ + 0_2^-$	4.0 $\times 10^{-30} \left(\frac{300}{T}\right)^{2.93}$	39
$0_4^+ + 0_2 \rightarrow 0_2^+ + 20_2$	1.3 $\times 10^{-4} \left(\frac{300}{T}\right)^{2.93} e^{-6000/T}$	b
$e + 0_4^+ \rightarrow 0 + 0 + 0_2$	$2.0 \times 10^{-6} (\frac{300}{T})$	33
$0_4^+ + 0 \rightarrow 0_2^+ + 0_3$	3.0×10^{-10}	34
$0_2^+ + 0_3^- + 0_2^- \rightarrow 0_5^+ + 0_2^-$	1.0 $\times 10^{-28} \left(\frac{300}{T}\right)^3$	С
$0_5^+ + 0_2 \rightarrow 0_2^+ + 0_3 + 0_2$	3.3 $\times 10^{-4} \left(\frac{300}{T}\right)^3 e^{-7850/T}$	d
$e + 0_5^+ \rightarrow 0 + 20_2$	$2.0 \times 10^{-6} (\frac{300}{T})$	е
$0_4^+ + 0_3 \rightarrow 0_5^+ + 0_2$	1.0 X 10 ⁻¹⁰	f
$0_5^+ + 0_2 \rightarrow 0_4^+ + 0_3$	1.0 X 10 ⁻¹¹ e ^{-1850/T}	f
$e + 0_2 + 0_2 \rightarrow 0_2 + 0_2$	1.4 $\times 10^{-29} \left(\frac{300}{T}\right)_{e}^{-600/T}$	33
$0_{2}^{-} + 0 \rightarrow e + 0_{3}$	1.5 X 10 ⁻¹⁰	33
$0_{2}^{-} + 0 \rightarrow 0^{-} + 0_{2}$	1.5 X 10 ⁻¹⁰	33
$0_2^- + 0_2^- (^1 \Delta_g^-) \rightarrow e + 20_2^-$	2.0 X 10 ⁻¹⁰	34

Table 2. Reactions and Rate Coefficients for the Pure Oxygen System (Continued).

Reaction	Rate Coefficient	Reference or Footnote
0 + 0 + e + 0 ₂	2.0 X 10 ⁻¹⁰	33
$0^{-} + 0_2 (^{1}\Delta_g) \rightarrow e + 0_3$	3.0 X 10 ⁻¹⁰	34
$0_{2}^{-} + 0_{3} \rightarrow 0_{3}^{-} + 0_{2}$	4.0 X 10 ⁻¹⁰	33
$0_3^- + 0 \rightarrow 0_2^- + 0_2$	2.5 X 10 ⁻¹⁰	34
$0_3^- + 0_3 \rightarrow 0_2^- + 20_2$	$3.0 \times 10^{-11} e^{-1000/T}$	g
$0_{2}^{-} + 0_{2} + 0_{2} \rightarrow 0_{4}^{-} + 0_{2}$	3.5 $\times 10^{-31} \left(\frac{300}{T}\right)$	33
$0_{4}^{-} + 0 \rightarrow 0_{3}^{-} + 0_{2}$	4.0 X 10 ⁻¹⁰	34
$0_4^- + 0_3 \rightarrow 0_3^- + 20_2$	3.0 X 10 ⁻¹⁰	33
$0^+ + 0_2^- \rightarrow 0 + 0_2$	2.0 X 10 ⁻⁶	h
$0^+ + 0_4^- \rightarrow 0 + 20_2$	2.0 X 10 ⁻⁶	h
$0_2^+ + 0^- \rightarrow 0 + 0_2$	2.0 X 10 ⁻⁶	h,i
$0_2^+ + 0_2^- \rightarrow 0 + 0 + 0_2$	2.0 X 10 ⁻⁶	h
$0_2^+ + 0_3^- \rightarrow 0_2 + 0_3$	2.0 X 10 ⁻⁶	h
$0_2^+ + 0_4^- \rightarrow 0 + 0 + 20_2$	2.0 X 10 ⁻⁶	h
$0_4^+ + 0^- \rightarrow 0 + 20_2$	2.0 X 10 ⁻⁶	h

Table 2. Reactions and Rate Coefficients for the Pure Oxygen System (Concluded).

Reaction	Rate Coefficient	Reference or Footnote
$0_4^+ + 0_2^- \rightarrow 0 + 0 + 20_2$	2.0 X 10 ⁻⁶	h
$0_4^+ + 0_3^- \rightarrow 0_3^- + 20_2^-$	2.0 X 10 ⁻⁶	h
$0_4^+ + 0_4^- \rightarrow 0 + 0 + 30_2$	2.0 X 10 ⁻⁶	h
$0_5^+ + 0_2^- \rightarrow 0_3^- + 20_2^-$	2.0 X 10 ⁻⁶	h
$0_5^+ + 0_3^- + 20_3 + 0_2$	2.0 X 10 ⁻⁶	h
$0_5^+ + 0_4^- \rightarrow 0_3^- + 30_2^-$	2.0 X 10 ⁻⁶	h
$0 + 0_2 + 0_2 + 0_3 + 0_2$	6.9 $\times 10^{-34} \left(\frac{300}{T}\right)^{1.25}$	35
$0 + 0_3 \rightarrow 0_2 + 0_2$	8.0 X 10 ⁻¹² e ^{-2060/T}	36
$0 + 0 + 0_2 \rightarrow 0_2 + 0_2$	4.8 $\times 10^{-33} \left(\frac{300}{T}\right)^2$	j
$0 (^{1}D) + 0_{2} + 0 + 0_{2} (^{1}\Sigma_{g}^{+})$	2.6 X 10 ⁻¹¹ e ^{67/T}	36
$0 (^{1}D) + 0_{2} \rightarrow 0 + 0_{2}$	6.0 X 10 ⁻¹² e ^{67/T}	37
$0 (^{1}D) + 0_{3} + 0_{2} + 0_{2}$	1.2 X 10 ⁻¹⁰	36
$0 (^{1}D) + 0_{3} + 0 + 0 + 0_{2}$	1.2 X 10 ⁻¹⁰	36
$0_2 (^{1}\Delta_g) + 0_2 + 0_2 + 0_2$	2.2 X $10^{-18} \left(\frac{T}{300}\right)^{0.8}$	38
$0_2 (^1 \Sigma_g^+) + 0_2 + 0_2 (^1 \Delta_g) + 0_2$	3.8 X 10-17	47,48
$0_2 (^{1}\Delta_g) + 0_3 + 0 + 20_2$	1.2 X 10 ⁻¹¹ e ^{-2400/T}	38
$0_2 (^1 \Sigma_g^+) + 0_3 \rightarrow 0 + 20_2$	2.4 X 10 ⁻¹¹	49

Footnotes to Table 2

- a. Rate constant units for bimolecular reactions are cm^3/sec , and for termolecular reactions, cm^6/sec .
- b. The reverse rate coefficient is calculated using the forward rate coefficient and the equilibrium constant data of Payzant et al. (Ref. 40) and Conway and Janik (Ref. 41), $K_{\rm eq} = 3.0 \times 10^{-26} \; {\rm e}^{6000/T} \; {\rm cm}^3$.
- c. The value for He as the third body is about 1 x 10^{-28} cm⁶/sec at 200 K (Ref. 34). The same value is used for 0_2 as the third body at 300 K. Further, a temperature dependence of $(300/T)^3$ is assumed, as in the 0_2^{\dagger} + 0_2 + 0_2 reaction (Ref. 39).
- d. Calculated from the reverse rate coefficient and the equilibrium constant derived from the equilibria for $0_2^+ + 0_2 + 0_2 \Rightarrow 0_2^+ + 0_2$ and $0_2^+ + 0_3 \Rightarrow 0_2^+ + 0_2$ (see footnotes b and f).
- e. Assumed to have the same value as for $e + 0^+_4$.
- f. The forward reaction is estimated from other fast ion-neutral switching reactions (Ref. 34); the reverse reaction is calculated from the equilibrium constant of Dotan et al. (Ref. 42), $K_{\rm eq}$ = 10 e^{1850/T}.
- g. This rate coefficient was selected to have a room temperature value of about 10⁻¹² cm³/sec, as deduced in oxygen radiolysis experiments (Ref. 43), and an activation energy of about 2 kcal/mole, as deduced from temperature-dependent studies (Ref. 44). The present analysis suggests that this value may be an upper limit, and a more likely value is about half as large (see section 4.4).
- h. Estimated high pressure value from References 45 and 46.
- i. It is possible that the products are 0 + 0 + 0, but this reaction appears to be unimportant both in radiolysis experiments and in the ambient or disturbed atmospheres.
- j. The rate coefficient listed corresponds to N_2 as the third body (Refs. 33 and 38). Insufficient data are available at temperatures below 500 K with O_2 as the collision partner.

The importance of the clusters is principally in their greater thermochemical stability, which might prevent them from breaking up to yield oxygen atoms when they recombine with oppositely-charged ions (see next section). Ion molecule interchange reactions can also lead to ions of greater stability:

$$0_4^+ + 0_3 \rightarrow 0_5^+ + 0_2$$
. (5)

Such reactions have not been considered in past radiolysis studies.

In the case of negative ions, molecules with large electron affinities can scavenge negative charge by charge transfer. Two such reactions that are important in oxygen radiolysis are

$$0_{2}^{-} + 0_{3} \rightarrow 0_{3}^{-} + 0_{2},$$
 (6)

and
$$0_4^- + 0_3^- + 0_2^- + 0_2^-$$
 (7)

Previous radiolysis analyses have omitted the latter reaction.

Certain reactions can detach electrons or reduce complex ions to simpler ones, for example:

$$0_2^- + 0 \rightarrow e + 0_3$$
, (8)

$$0_3^- + 0 + 0_2^- + 0_2,$$
 (9)

$$0_4^- + 0 + 0_3^- + 0_2^-$$
 (10)

However, at the high pressures typical of radiolysis experiments, such recycling reactions are overwhelmed by clustering reactions.

At pressures near one atmosphere, larger cluster ions such as 0_6^+ and 0_6^- are known to form, but few of their reaction rate coefficients are known. Equilibrium data (Ref. 41) indicate that their concentrations are small except below $200^{\,0}\mathrm{K}$, so they have not been included in the present analysis.

3.3 ION-ION RECOMBINATION.

In radiolysis experiments, the charge produced by irradiation is neutralized primarily by recombination of positive with negative ions. In the present case, the positive ions are predominantly 0_2^+ and 0_4^+ , and the recombination products can be 0_2 or, if enough energy is available for dissociation, 0+0. The distinction is important because the 0 atoms generated by ion-ion recombination can eventually form ozone and increase the ozone yield. If all recombination occurs via stable ions without dissociation, no net ozone production from ions results.

In the most recently published discussion of the ozone yield under various conditions of oxygen radiolysis (Ref. 1), it was suggested that reactions of the following sort lead to 0-atom generation:

$$\begin{pmatrix} 0_2^+ \text{ or } 0_4^+ \end{pmatrix} + \begin{pmatrix} 0_2^- \text{ or } 0_4^- \end{pmatrix} \rightarrow 0 + 0 + n 0_2,$$
 (11)

while the following reactions do not:

$$\left(0_{2}^{+} \text{ or } 0_{4}^{+}\right) + 0_{3}^{-} \rightarrow 0_{3} + n 0_{2}.$$
 (12)

These suggestions have been adopted in the present calculations. We also assume that the recombination of 0_5^+ with negative ions leads to $0_2^- + 0_3^-$ without 0-atom production.

Measurements of ion-ion recombination at atmospheric pressure suggest a typical recombination coefficient of about 2 x 10^{-6} cm³/sec (Refs. 45, 46). A specific study has been made by Wood et al. (Ref. 50) of ion-ion recombination rates in oxygen subject to low level continuous gamma irradiation. At pressures of 100-300 torr, an average coefficient of about 1.1 x 10^{-6} cm³/sec was measured. At a pressure around 30 torr the coefficient decreased to about 0.9 x 10^{-6} cm³/sec. Since most radiolysis experiments have been conducted at pressures near one atmosphere, in the present calculations a single ion-ion recombination coefficient of 2 x 10^{-6} cm³/sec has been adopted for all pairs of ions at all pressures and temperatures.

3.4 OZONE CHEMISTRY.

The principal neutral reactions controlling the ozone concentration are the production mechanism:

$$0 + 0_2 + 0_2 \rightarrow 0_3 + 0_2,$$
 (13)

and the destruction mechanism:

$$0 + 0_3 \rightarrow 0_2 + 0_2$$
. (14)

Such reactions can occur for both ground and excited-state oxygen atoms, with different rate coefficients. In the radiolysis experiments of interest, excited oxygen atoms are so rapidly quenched by oxygen molecules that only the ground-state reactions need to be considered.

Ozone can also be destroyed by ion-molecule reactions, such as

$$0_{2}^{-} + 0_{3} \rightarrow 0_{3}^{-} + 0_{2}$$
, (6)

$$O_3^- + O_3 \rightarrow O_2^- + 2O_2^-$$
 (15)

This reaction sequence is "catalytic" in that the second reaction regenerates the $0\frac{1}{2}$ lost in the first reaction. A similar pair of destruction reactions could involve $0\frac{1}{4}$ and $0\frac{1}{5}$ in place of $0\frac{1}{2}$ and $0\frac{1}{3}$, respectively.

The excited molecular oxygen states $O_2(^1\Delta)$ and $O_2(^1\Sigma)$ can dissociate ozone:

$$O_2(^1\Delta) + O_3 \rightarrow O + 2 O_2,$$
 (16)

$$O_2({}^{1}\Sigma) + O_3 \rightarrow O + 2 O_2.$$
 (17)

However, at the pressures typical of radiolysis experiments, almost all of the oxygen atoms thus produced quickly react to re-form ozone. Hence, the role of these excited states is minor.

An upper limit to the amount of ozone that can be produced by radiolysis may be derived by neglecting all ozone destruction processes. With this neglect, every oxygen atom that is created yields one ozone molecule. From table 1, for each 100 eV of energy deposited in O_2 , 7.0 oxygen atoms are produced, and also 1.0 O^+ ions, which rapidly charge exchange with O_2 to yield another oxygen atom. Hence, from these processes alone, $G(O_3) = 8.0$, where the G value is defined as the number of particles produced per 100 eV deposited. In addition, as discussed in section 3.2, each positive ion

created will, when it recombines with an electron or negative ion, produce two oxygen atoms, provided that other reactions have not converted the ions to more stable species like 0^-_3 . Since G (positive ions) = 3.25 (table 1), ion recombination can increase $G(0_3)$ by up to 6.5, giving a total upper limit of $G(0_3)$ = 8.0 + 6.5 = 14.5.

3.5 CHEMISTRY COMPUTATION METHOD.

In the present study, time-dependent chemistry calculations have been made for situations corresponding to published oxygen radiolysis experiments. These calculations used the CHEM computer code developed at RDA, which employs a Gear integrator (Ref. 51) to solve numerically the coupled differential equations for the chemical kinetics. The calculations employed the initial production rates listed in table 1 and the rate coefficients listed in table 2. The results will be discussed in the next section.

SECTION 4

COMPUTED RESULTS AND COMPARISON WITH EXPERIMENT

4.1 OZONE PRODUCTION IN PULSE RADIOLYSIS.

Several studies have been carried out in which intense short-duration pulses of electrons were used to irradiate oxygen (Refs. 1, 52-54). The electron pulse generators used had peak electron energies of 600 keV to 2 MeV and pulse half-widths of around 30 nsec.

Typical species concentrations calculated for the pulse radiolysis of oxygen at 1 atmosphere are shown in Figure 1. The principal ions are 0_4^+ , 0_2^- and 0_4^- . The concentrations of electrons and 0_2^+ are lower due to efficient attachment and clustering reactions. The ionic charge builds up during the period of electron irradiation, reaches a maximum, and then decays by ion-ion recombination, approximately as t^{-1} after the end of the pulse. At late times, as ozone appears, the ions 0_3^- and 0_5^+ become appreciable; however, by this time most of the ionization has dissipated.

Atomic oxygen, which is produced directly by electron bombardment, accumulates throughout the irradiation period. The deionization processes contribute additional atomic oxygen, so 0 continues to build up early in the deionization phase. At later times substantial ozone concentrations are produced from the association of 0 with 0_2 . Most of the ionization has disappeared by this time. The molecular oxygen metastable states $^1\!\Delta$ and $^1\!\Sigma$ decay over still longer time scales, and their dissociating reactions with 0_3 tend to maintain a

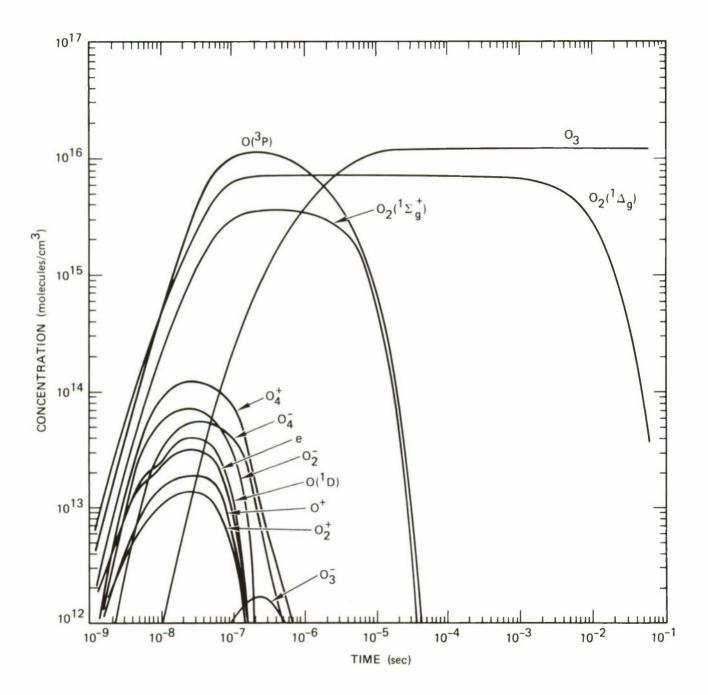


Figure 1. Time-dependent species concentrations in oxygen subject to a brief irradiation pulse varying with time as t $\exp(-t/t_0)$, where $t_0 = 22$ nsec. The integrated dose is 1 x 10⁶ rads, pressure 760 torr, and temperature 293 K.

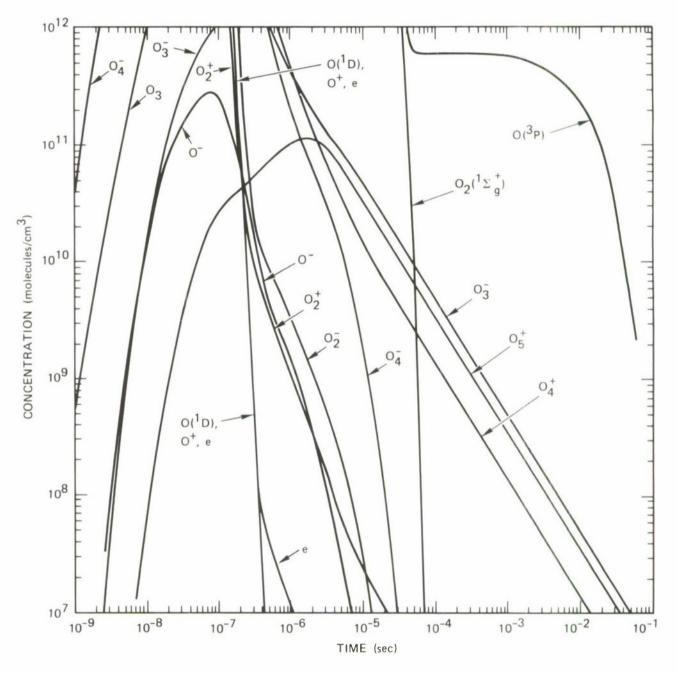


Figure 1. Time-dependent species concentrations in oxygen subject to a brief irradiation pulse varying with time as t $\exp(-t/t_0)$, where $t_0 = 22$ nsec. The integrated dose is 1 x 10^6 rads, pressure 760 torr, and temperature 293 K. concluded).

low-level concentration of atomic oxygen, but this has little effect on the $\mathbf{0}_3$ concentration or yield.

In table 3, predicted ozone G values from pulse radiolysis are compared with measured values. The agreement is reasonably good, considering the uncertainties in the computations and The principal source of uncertainty in the the measurements. calculations is probably the initial yields of various species in oxygen under electron bombardment. All of the pulse radiolysis experiments, and the corresponding calculations, show a high ozone yield, near the theoretical limit. Since figure 1 shows that at atmospheric pressure most of the charged-particle recombination occurs between 0_4^+ , 0_4^- and 0_2^- , this high ozone yield implies that recombination of these ions yields two oxygen atoms, as assumed in Table 2. Although these reactions are not listed in the DNA Reaction Rate Handbook (Ref. 33), the similar reaction, $0_2^+ + 0_2^- \rightarrow 0_2^- + 0_2^-$, is included. The present results strongly suggest that the products should be revised to $0_2 + 2 0$.

Similar chemistry calculations were made for pulse radiolysis at 0.1 atm. The results show that 0^+_2 , 0^+_4 and e are the major charged species in this case. The high ozone yield observed at this pressure by Willis et al. (Ref. 1) then indicates that e + 0^+_4 yields two oxygen atoms. This result has been generally assumed (Ref. 33) but not previously confirmed.

4.2 EFFECT OF ELECTRON SCAVENGERS ON OZONE YIELD IN PULSE RADIOLYSIS.

An electron scavenger like sulfur hexafluoride (SF $_6$) can attach electrons to form SF $_6$, a negative ion that is more stable than 0^-_2 or 0^-_4 , and hence may not produce oxygen atoms

Table 3. Ozone yield in the pulse radiolysis of oxygen. High dose rates: 10^{11} - 10^{13} rad/s (1 rad = 100 erg/g). Gas temperature: 293 - 298 K.

Type of determination	O ₂ pressure (torr)	G(O ₃) molec/100 eV)	Reference
Measurement	560	13.0*	52
Measurement	760	13.8±0.7	54
Measurement	30 800	12.8 <u>±</u> 0.6	1
Calculation	83, 830	14.0±0.9	1
Calculation	76, 760	13.8-14.4	This work
Theoretical limit	and the	14.5	This work

 $^{^{*}\}mathrm{G}$ value corrected for dosimetry in Ref. 1.

when it recombines with 0_2^+ or 0_4^+ . These and other relevant reactions of ${\rm SF}_6$ are summarized in table 4. Using these reactions and rate constants, chemistry calculations were made for the pulse radiolysis of oxygen containing 1 percent (by volume) of ${\rm SF}_6$, corresponding to three published experiments (Refs. 1, 52, 53).

Table 5 compares the measured and calculated ozone yields. The calculated $G(O_3)=8.0-8.1$ (varying slightly with pressure) lies near the center of the range of measured values, $G(O_3)=6.3-9.0$. Since the calculations indicate that the major recombining ions are SF_6^- , O_4^+ and O_2^+ , the reasonable agreement of the calculated with the measured ozone yields confirms the assumption (table 4) that recombination of these ions does not yield oxygen atoms.

4.3 OZONE BUILDUP DURING CONTINUOUS RADIOLYSIS.

A number of researchers have studied the effects of continuous radiolysis of pure oxygen (Refs. 1, 44, 55, 56). Most of these studies used gamma ray sources and irradiation intensities up to about 10^7 rad/hr, with total doses exceeding 10^8 rads in some cases. The difference between the energy sources employed in the pulse and steady radiolysis experiments (electrons versus gamma rays) has no direct bearing on the ion chemistry or ozone yield; the primary ionization products per unit energy absorbed are essentially the same in both cases. However, the steady radiolysis experiments involved much longer irradiation times and therefore significant ozone accumulation during the period of irradiation. Consequently, ions like 0^-_3 and 0^+_5 are expected to be more important here than in the pulse radiolysis experiments.

Table 4. SF₆ chemistry.

Reaction	Rate Coefficient (a)	Reference or Footnote
$e + SF_6 \rightarrow SF_6^-$	2.2 x 10 ⁻⁷	57, 58
$SF_6 + O_2^- \rightarrow SF_6^- + O_2$	7.0 x 10 ⁻¹¹	34
$SF_6 + O_4^- \rightarrow SF_6^- + 2 O_2$	7.0 x 10 ⁻¹¹	ь
$SF_6^- + O \rightarrow O^- + SF_6$	5.0 x 10 ⁻¹¹	34
$SF_6^- + O_3 \rightarrow O_3^- + SF_6$	3.2 x 10 ⁻¹¹	34
SF ₆ + O ⁺ → SF ₆ + O	2.0 x 10 ⁻⁶	ь
$SF_6^- + O_2^+ \rightarrow SF_6 + O_2$	2.0 x 10 ⁻⁶	ь
$SF_6^- + O_4^+ \rightarrow SF_6 + 2 O_2$	2.0 x 10 ⁻⁶	ь
$SF_6^- + O_5^+ + SF_6 + O_2 + O_3$	2.0 x 10 ⁻⁶	р

Footnotes:

- a. All rate coefficients have units of cm^3/sec .
- b. Assumed by analogy.

Table 5. Ozone yield in pulse radiolysis of oxygen containing 1 percent (by volume) of sulfur hexafluoride; temperatures near 295 K.

Pressure (torr)	G(O ₃)	References
560	9.1*	52
200-2000	7.9†	53
400, 700	6.3±0.6	1
76, 760	8.0,8.1	This work
	8.0	This work
	560 200-2000 400, 700	560 9.1* 200-2000 7.9† 400, 700 6.3±0.6 76, 760 8.0,8.1

^{*}G value corrected for dosimetry as in Ref. 1.

[†]Normalized using $G(O_3) = 12.8$ in pure oxygen at 760 torr (see table 3).

The present calculations for continuous radiolysis situations bear out this expectation. Although, during the very earliest period of the irradiation (corresponding to integrated doses of less than 10^3 rads) the ozone concentration is so small that little 0^-_3 or 0^+_5 is predicted to form, by the time the ozone accumulates to levels that are easily measureable (around 10^4 rads) 0^-_3 becomes the dominant negative ion, and 0^+_4 and 0^+_5 the dominant positive ions. Consequently, the reasonable agreement of the calculated and measured ozone yields shown in table 6 supports the assumption (table 2) that recombination of 0^-_3 with 0^+_4 or 0^+_5 does not lead to dissociation. (The discrepancy of about 25 percent between the calculations and the measurements is not readily explainable but probably has little practical significance.)

As the oxygen irradiation time is lengthened and the total dose increases above about 10^5 rads, the calculated ozone production rate slows, due primarily to the assumed process, $0^-_3 + 0^-_3 + 0^-_2 + 2 \cdot 0^-_2$ (Ref. 43), which effectively destroys two ozone molecules per reaction. (The well-known ozone destruction reaction, $0 + 0^-_3 + 2 \cdot 0^-_2$, is much slower under most radiolysis conditions.) A qualitatively similar behavior has been observed in two sets of high-dose radiolysis measurements (Ref. 44, 56). However, Kircher et al. (Ref. 44) used "commercial" oxygen containing impurities that may have affected the ozone chemistry; moreover, their pressures were much higher than atmospheric. Consequently, no comparisons of Kircher's measurements with chemical calculations have been attempted.

Sears and Sutherland (Ref. 56) carried out experiments near atmospheric pressure using both all-glass vessels and quartz vessels having stopcocks lubricated with fluorocarbon grease.

Table 6. "Initial" ozone yield in the steady radiolysis of oxygen. ("Initial" implies low total dose, ~10 rads.)
Gas temperatures are near 295 K unless otherwise indicated.

Type of determination	⁰ ₂ pressure	G(0 ₃)	References
Measurement	(2-10)x10 ⁴	5.9*	44
Measurement (200 K)	760	6.7* †	5 5
Measurement	665	6†‡	56
Measurement (77 K)	not stated	6.2 <u>±</u> 0.6	1
Theoretical estimate	and non	7.5±0.9	1
Calculation	665	8 - 9	This work
Theoretical limit, without ionic contri	 b.	8.0	This work

^{*}G value corrected for ozone-iodide stoichiometry in Ref. 1. †G value corrected for dosimetry in Ref. 1. ‡See Ref. 1.

They observed that the ozone yields were greater in the vessels with stopcocks. This observation is consistent with the fact that many fluorine compounds will remove the negative charge from ions like 0_3^- , and hence interfere with the $0_3^- + 0_3^-$ destruction mechanism.

Unfortunately, Sears and Sutherland reported dose-dependent measurements only for the vessels with greased stopcocks. Their results are plotted in figure 2, along with calculations for pure oxygen and oxygen contaminated with 0.0044 percent (44 ppmv) of a hypothetical electron scavenger. The scavenger is assumed to have the same reactions as SF₆ (table 4), except that the charge transfer reactions with 0 and 0^-_3 proceed in the opposite direction (i.e., the scavenger has a greater electron affinity than 0 and 0^-_3). The experimental points in figure 2 fall between the calculated points for pure and contaminated oxygen, strongly suggesting that the stopcock grease is producing an electron scavenger that removes the charge from 0^-_3

4.4 STEADY-STATE OZONE CONCENTRATIONS IN CONTINUOUS RADIOLYSIS.

With long enough irradiation times (corresponding to doses above approximately 10⁷ rads) the chemical processes reach a steady state where ozone is destroyed as fast as it is formed. The ozone concentration thus eventually approaches a constant value. Sears and Sutherland (Ref. 56) present measurements for this concentration in all-glass (uncontaminated) vessels, as a function of the dose rate. These values are compared in figure 3 with values calculated using the reaction rates in table 2. The agreement is fairly good. The experimental values vary approximately—and the calculated values almost exactly—with the square root of

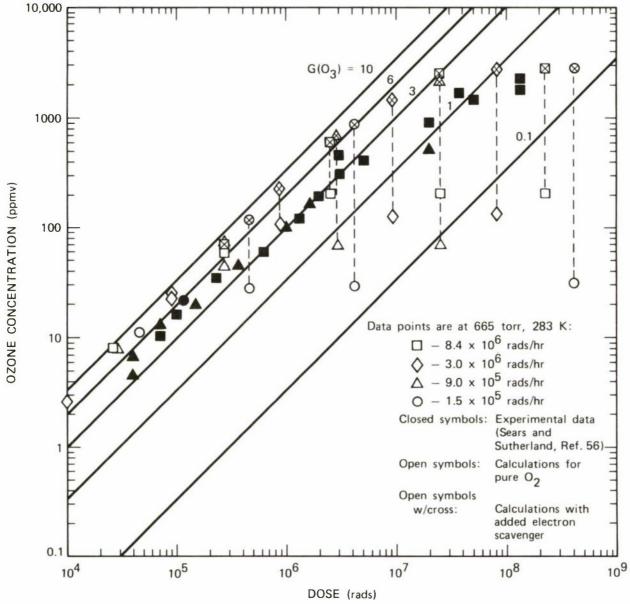


Figure 2. Measured and calculated ozone concentrations versus cumulative radiation dose for the continuous (slow) radiolysis of oxygen at 665 torr and 283 K. The diagonal lines correspond to loci of constant G values. For the calculated points connected by vertical dashed lines, the lower values correspond to pure oxygen while the upper values correspond to oxygen contaminated with 44 ppmv of a hypothetical electron scavenger (see text).

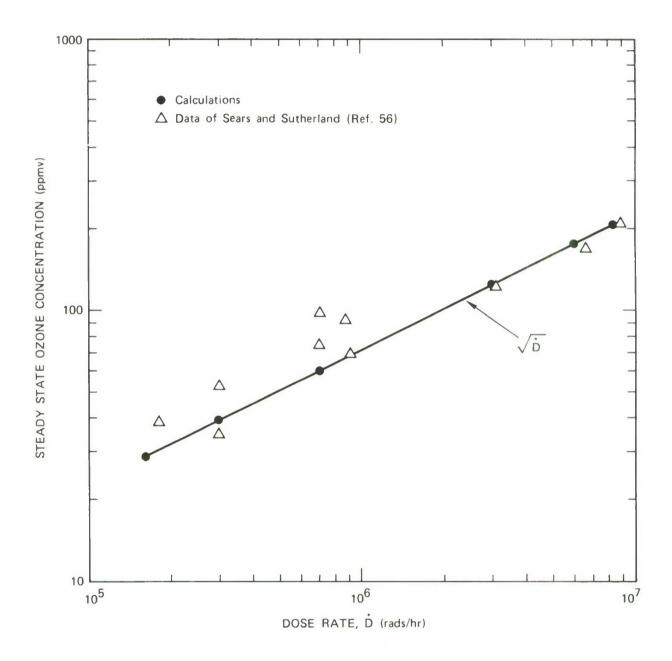


Figure 3. Comparison of measured and calculated steady-state ozone concentrations in irradiated oxygen as a function of dose rate.

the dose rate (solid line on figure 3). As pointed out by Sears and Sutherland, this behavior expected when the primary removal process is the reaction of a major ion with $\mathbf{0}_3$, since in a steady state dominated by ion-ion recombination the ion concentration (of either sign) is proportional to the square root of the dose rate.

The agreement of the steady-state measurements with the calculations tends to support the assumed rate coefficient and products of the reaction $0_3^- + 0_3 \rightarrow 0_2^- + 2 0_2$ (table 2). However, before this support can be considered strong, alternative destruction reactions must be ruled out. Note first, that to fit the radiolysis observations, the species reacting with and destroying ozone must be an odd oxygen species, like O_{3}^{-} , O_{5}^{+} or $\mathrm{O(^{1}D)}$; excited O_{2} molecules can only break up ozone to give 0 + 02, where the 0 atoms immediately re-form ozone. The decrease in ozone destruction in the presence of fluorocarbon grease (figure 2) strongly suggests a negative ion, since many fluorine compounds have large electron affinities and can scavenge charge from other negative ions. Moreover, if a neutral species (other than ozone) were the reactant, its primary removal process would have to be reaction with itself or with another free radical or ion, in order to yield the observed square root behavior. Enough is known about the potentially significant neutral species (in fact, they have been included in table 2 and in the calculations) to show that none satisfy this requirement.

Possible ozone destruction by ions like 0^- , 0^+_3 and 0^+_5 should also be considered. However, the calculations show that the concentration of 0^- in the radiolysis experiments is always much too small to contribute significantly to the ozone loss. The 0^+_3 concentration is also small, because it is formed only

from 0^+ , which has a small concentration, and it is rapidly destroyed by charge transfer to 0_2 . (For these reasons, 0_3^+ was not included in the present chemical calculations.)

On the other hand, the concentration of 0_5^+ is large enough for it to destroy significant amounts of ozone, if the rate coefficient were large enough. At the lowest and highest dose rates employed by Sears and Sutherland, 0_5^+ makes up about 20 and 60 percent, respectively, of the total positive ions, the remainder being 0_4^+ , which cannot destroy ozone. Hence, if the $0_5^+ + 0_3$ and $0_3^- + 0_3$ coefficients were equal, and both were a factor 1/1.6 smaller than the value used for the latter reaction in the calculations shown in figure 3, the low end of the curve would be raised about a factor of 1.6/1.2, while the high end would remain unchanged (neglecting the fairly small effect of the changed ozone amount on the $0_5^+/0_4^+$ ratio). This modified prediction would still fit the measurements reasonably well. However, if the 0_5^+ rate coefficient were twice the 0_3^- coefficient, there would be a significant discrepancy.

It follows that the $0_5^+ + 0_3^-$ rate coefficient is probably less than 6 x 10^{-13} cm $^3/s$ and may be much smaller. In addition, the value used for the 0_3^- reaction in the calculations 1 x 10^{-12} cm $^3/s$, is probably an upper limit, since it would have to be lowered if 0_5^+ contributes significantly to the destruction, or if the ion-ion recombination coefficient is smaller than the value of 2 x 10^{-6} cm $^3/s$ used in the calculations (compare Ref. 50), since this would increase the 0_3^- concentration. Accordingly, a reasonable choice for the room-temperature rate coefficient of the $0_3^- + 0_3^-$ reaction is 5 x 10^{-13} cm $^3/s$, which may be correct to within a factor of 2. It probably varies with temperature roughly as $\exp(-1000/T)$, as deduced by Kircher et al. (Ref. 44)

for the temperature dependence of ozone destruction in high-pressure oxygen radiolysis. However, since the temperature may also shift the relative ion concentrations and change the ionion recombination coefficient, the observed temperature dependence may not be attributable solely to the $0\frac{\pi}{3} + 0\frac{\pi}{3}$ reaction.

SECTION 5

CONCLUSIONS

The detailed electron-impact and chemical calculations carried out in this study have given results which agree within 25 percent with published measurements of ozone production in a variety of radiolysis experiments. The calculations employed electron-impact cross sections newly deduced from published data, and standard sets of reaction rate coefficients, with the addition of the following ion-ion recombination reactions:

$$0_4^+ + 0_2^- \rightarrow 2 \ 0 + 2 \ 0_2^-,$$
 (18)

$$0_{4}^{+} + 0_{4}^{-} \rightarrow 2 \ 0 + 3 \ 0_{2}^{-},$$
 (19)

$$0_4^+ + 0_3^- \rightarrow 2 \ 0_2 + 0_3,$$
 (20)

$$0_5^+ + 0_3^- \rightarrow 0_2^- + 2 0_3^-.$$
 (21)

(Some other recombination reactions were also added for completeness, but they have a negligible effect on the results.)

In conventional ion recombination experiments it has not been possible to identify the products of these reactions. The present analysis of the radiolysis experiments permits confident assignment of these products (except that in the last two reactions the $\mathbf{0}_3$ might be replaced by $\mathbf{0} + \mathbf{0}_2$). Moreover, the products of the first reaction strongly imply that the $\mathbf{0}_2^+$ + $\mathbf{0}_2^-$ recombination reaction similarly yields

 2 0 $^+$ 0 2, instead of 2 0 2 as conventionally assumed. In addition, the analysis of the steady-state ozone concentration attained by prolonged irradiation, as a function of the dose rate, suggests that the ozone-destroying reaction

$$0_3^- + 0_3 + 0_2^- + 2 0_2$$
 (15)

occurs with a rate coefficient within a factor of two of $5 \times 10^{-13} \text{ cm}^3/\text{s}$ at room temperature, and probably with a temperature variation of about exp (-1000/T). The reaction,

$$0_5^+ + 0_3 \rightarrow 0_4^+ + 2 0_2$$
 (22)

may also provide a significant ozone sink. However, it cannot be the dominant loss mechanism, so its rate constant cannot be larger than about $6 \times 10^{-13} \text{ cm}^3/\text{s}$.

Calculations for oxygen containing traces of electron scavengers (such as fluorine compounds) show, in agreement with experiment, that these scavengers can affect the ozone yield in two opposing ways: In pulse radiolysis they can reduce the ozone yield by changing the ions to species that produce less atomic oxygen when they recombine. In continuous irradiation, where the ozone builds up and itself reduces the ozone yield by electron scavenging, addition of a stronger scavenger can remove the charge from 0^-_3 , thus inhibiting the 0^-_3 + 0^-_3 reaction that destroys ozone, and increasing the ozone yield.

SECTION 6

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